

SuperNOvA

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Assumptions

- Totally active detector (TASD)
- Detection of Supernova signals will require high live time front-end and DAQ
 - Online sparsification and Offline trigger
 - Won't be a free upgrade
- Possible in principle
 - Harvard prototyping a test system.
- Further challenges to come:
 - Backgrounds
 - DAQ rates

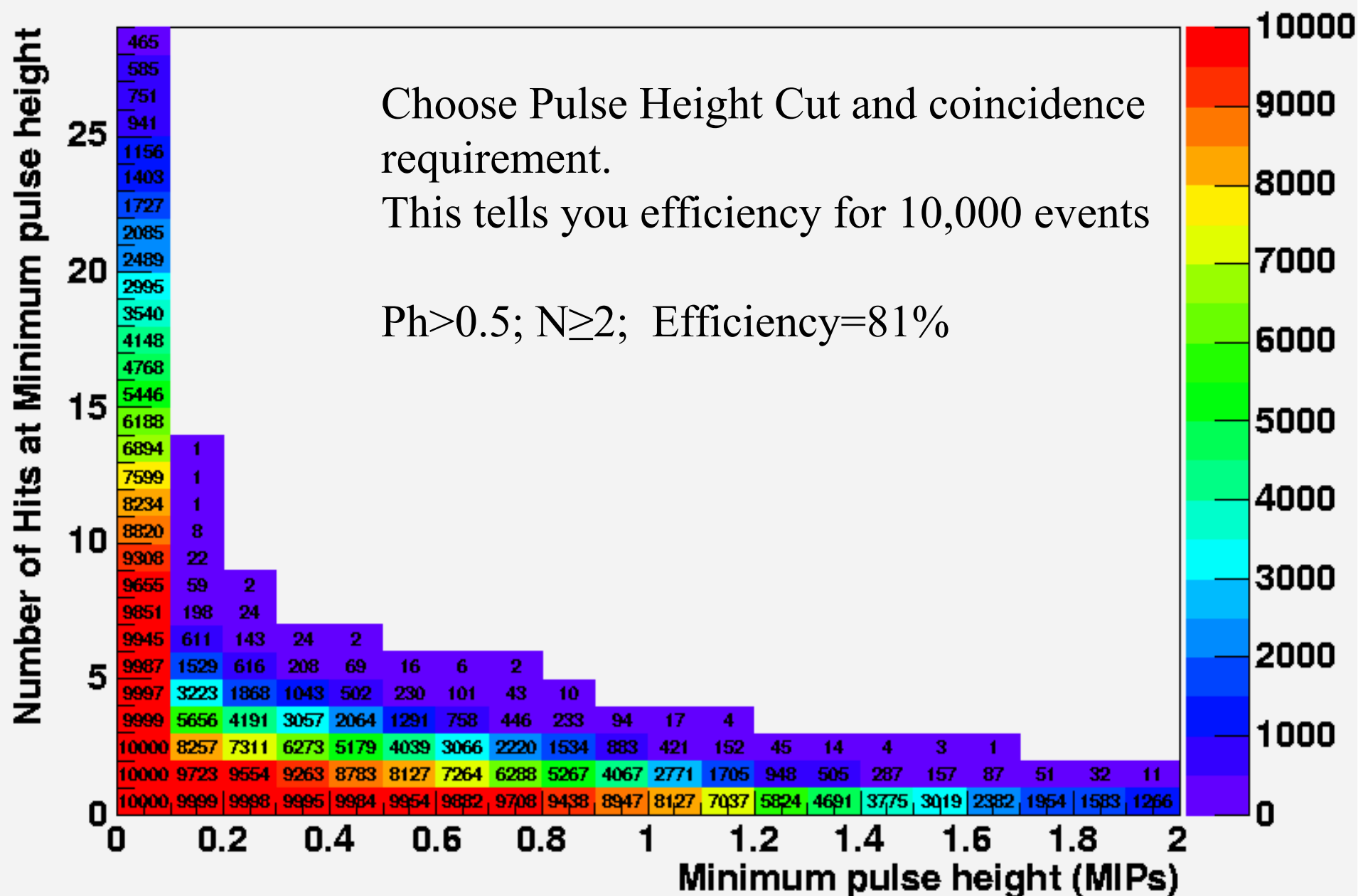
SuperNOvA Signals

- Primary Signal from $\bar{\nu}_e + p \rightarrow e^+ + n$
- Total signal –scaled from MiniBooNE
 - 8000 total interactions over ~ 10 s;
 - 4000 in first second
 - Energy peaks at 20MeV, falling to about 60MeV
 - Positron dumps this energy locally, a few hits.

SN Backgrounds

• Natural radioactivity	Small
• Cosmic Ray Muons	500kHz
• EM from CR EAS	300kHz
• Neutrons	30kHz
• Total	~MHz

SN signal Pulse Height Frequency



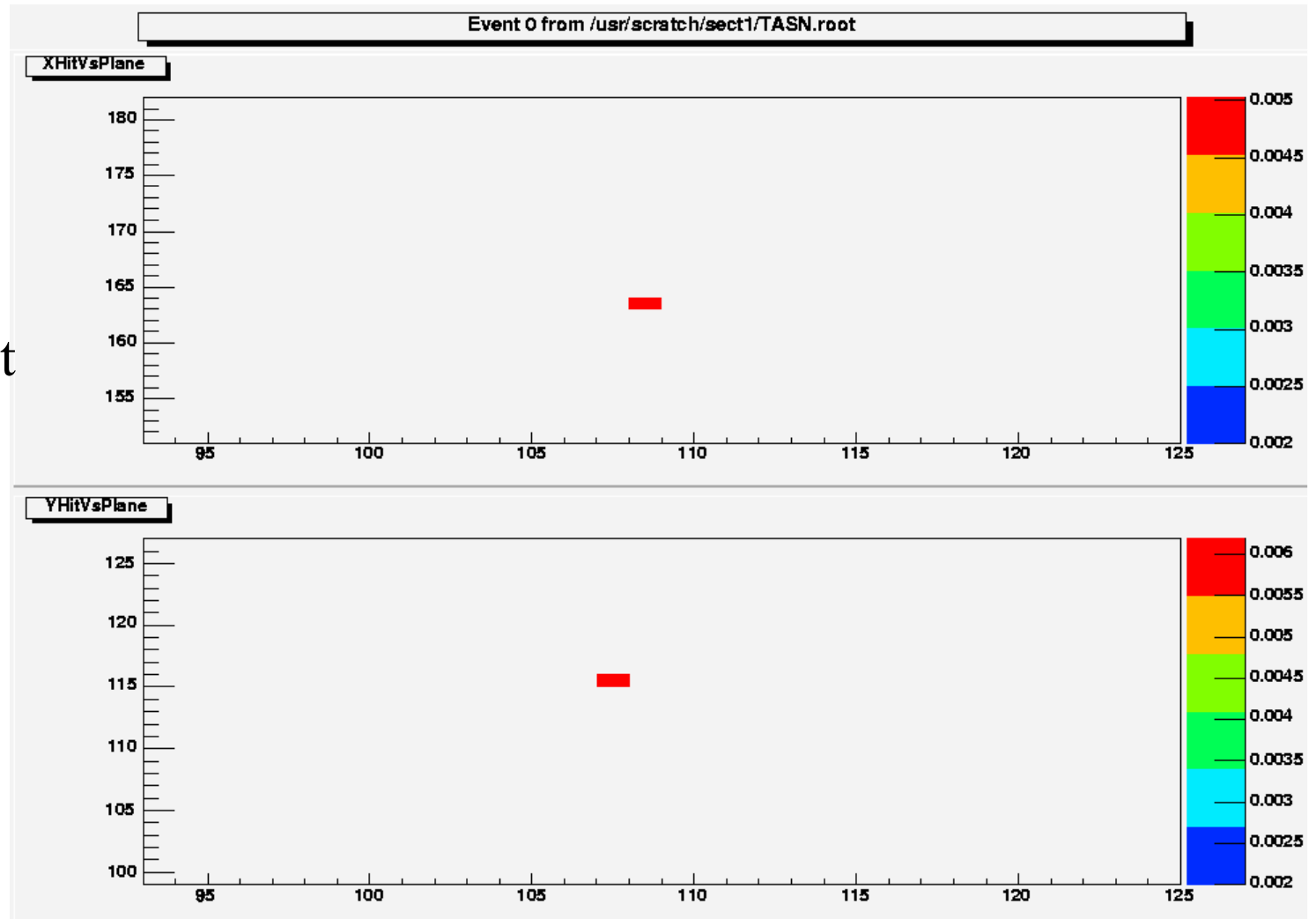
Signal With Reality

- Out of 8000, concentrate on 4000 from first 1 second, best possible S/N
- 80% trigger efficiency
 - 3200 events in 1 second
- 50% reconstruction efficiency
 - 1600 events in 1 second

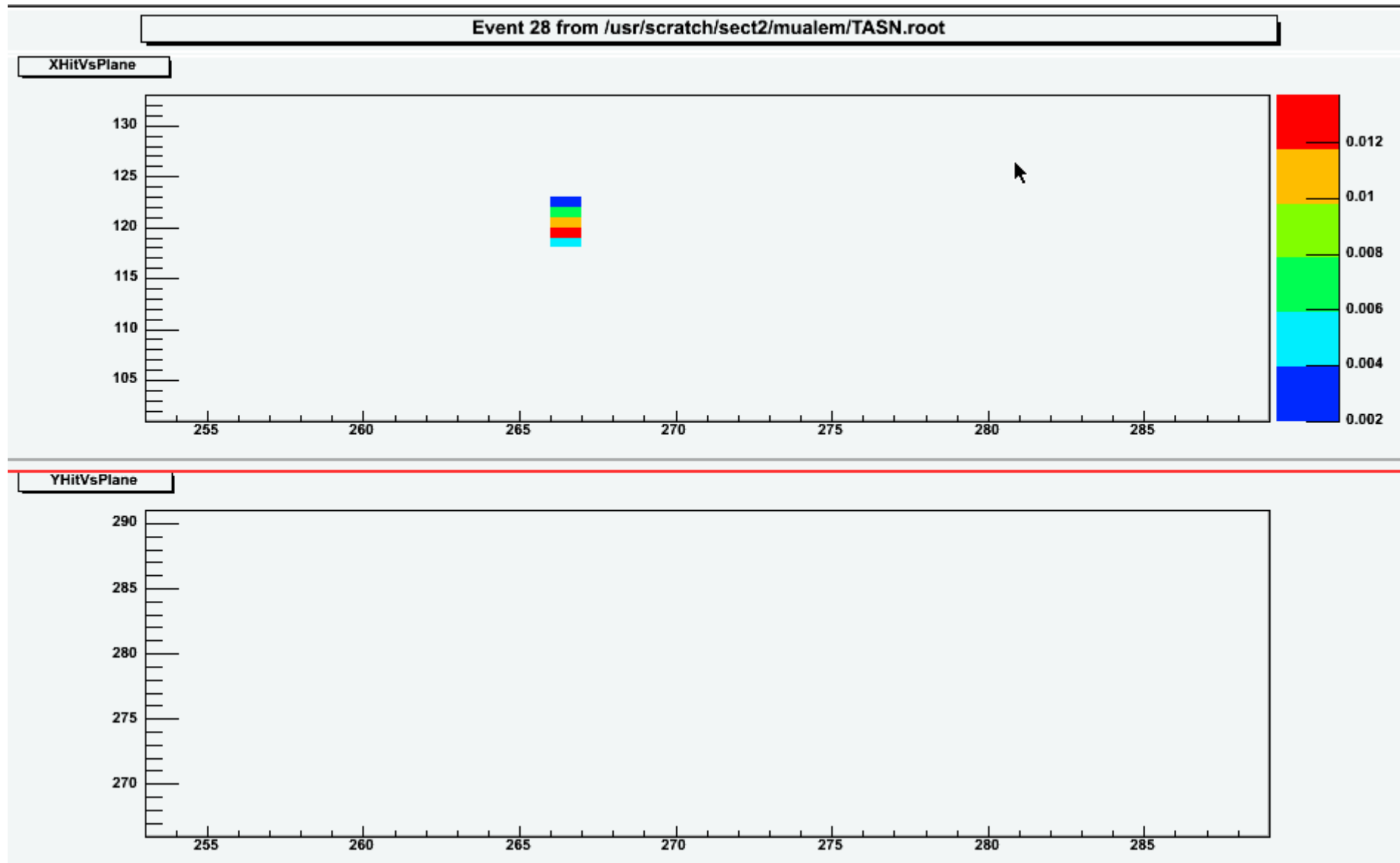
SN event passes trigger; accepted

16MeV e^+

One ~ 1 MIP hit
in each view



SN event passes trigger, gets cut anyway



Natural Radioactivity

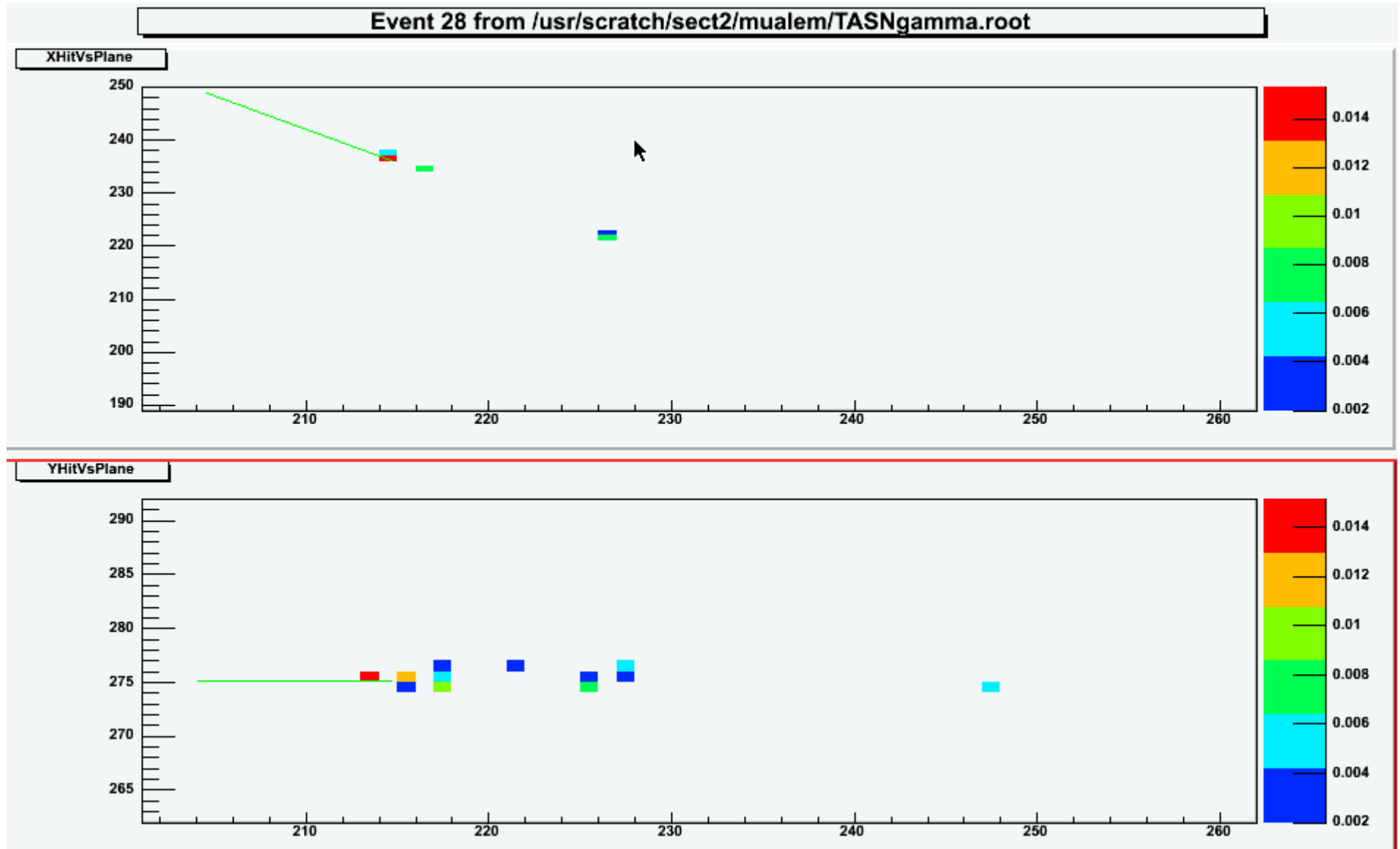
- Maximum energies around 2MeV, typical energies MUCH lower. Sparsification will get rid of vast majority of them.
- Setting threshold to coincident neighbors planes of $\frac{1}{2}$ far-end is equivalent to a threshold of 2MeV, strongly suppressing background at (offline) trigger level.
- Requiring reconstructed (atten corrected) energy of 8MeV effectively eliminates this background.
- Most betas too low in energy, and most gammas Compton scatter giving small fractions of their energy; these get sparsified away.
- Unclear what it will be precisely, K-40 from concrete, U, Th...
- MINOS NOT a good guide,
 - sensitive to 1/3pe $\sim 60\text{keV}$ – NOvA sensitive to 2MeV
 - Only bottom of detector near concrete...

EM background

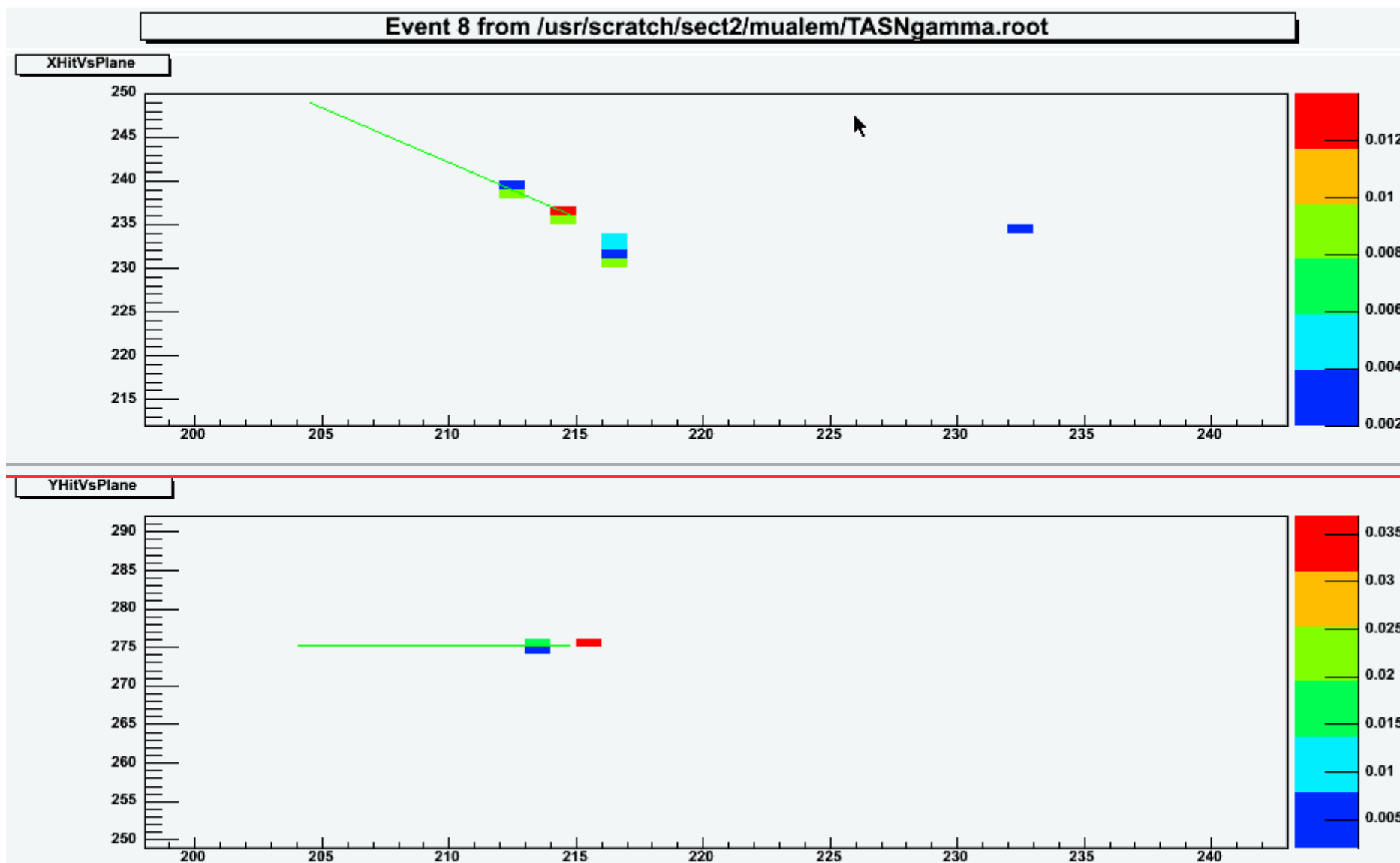
- Gammas and electrons from EAS
- Rate at surface 300kHz $E > 10\text{MeV}$
 - 150kHz $E > 100\text{MeV}$

CR gamma background

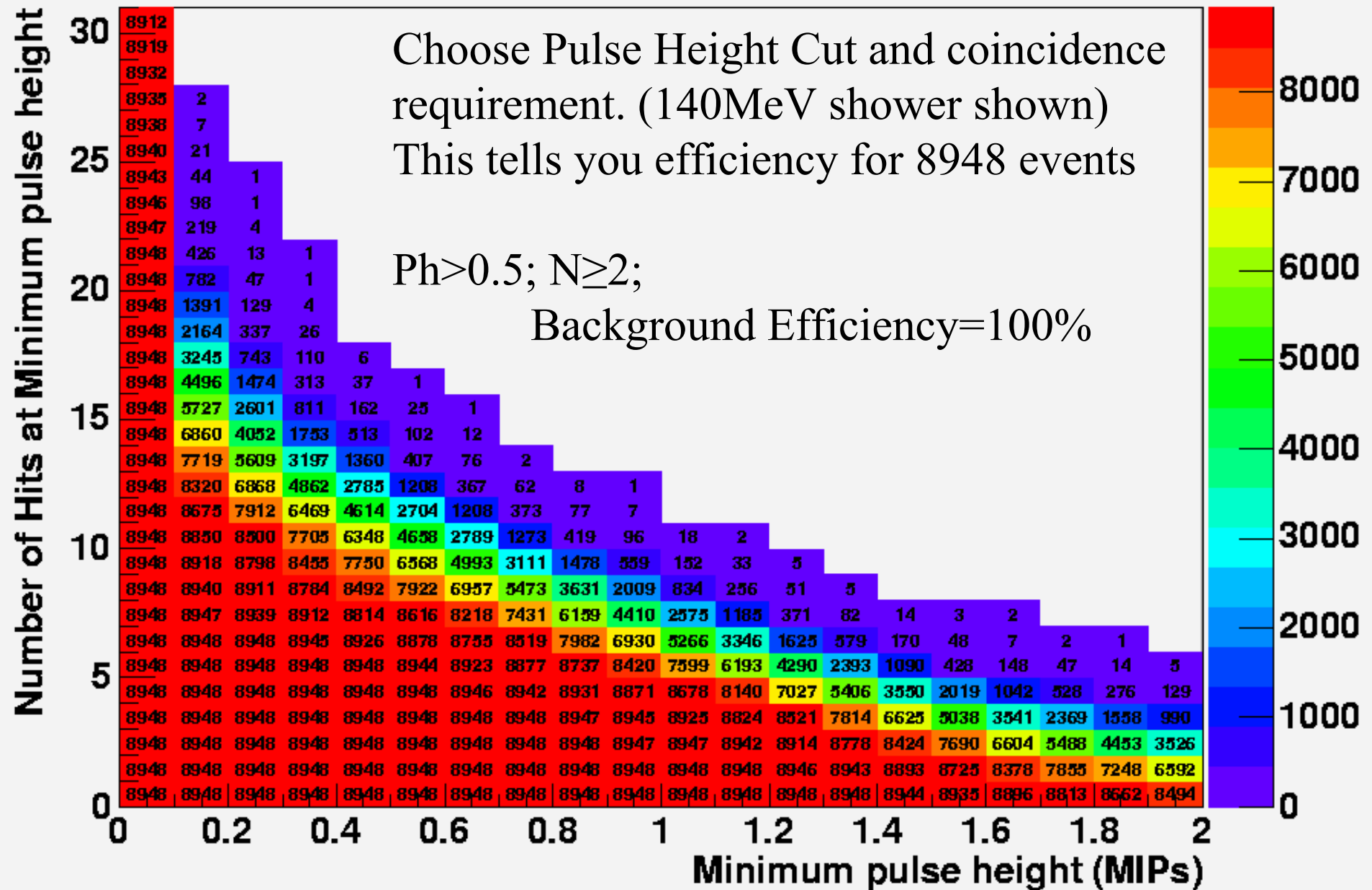
- Many vetoed, others eliminated by Etot



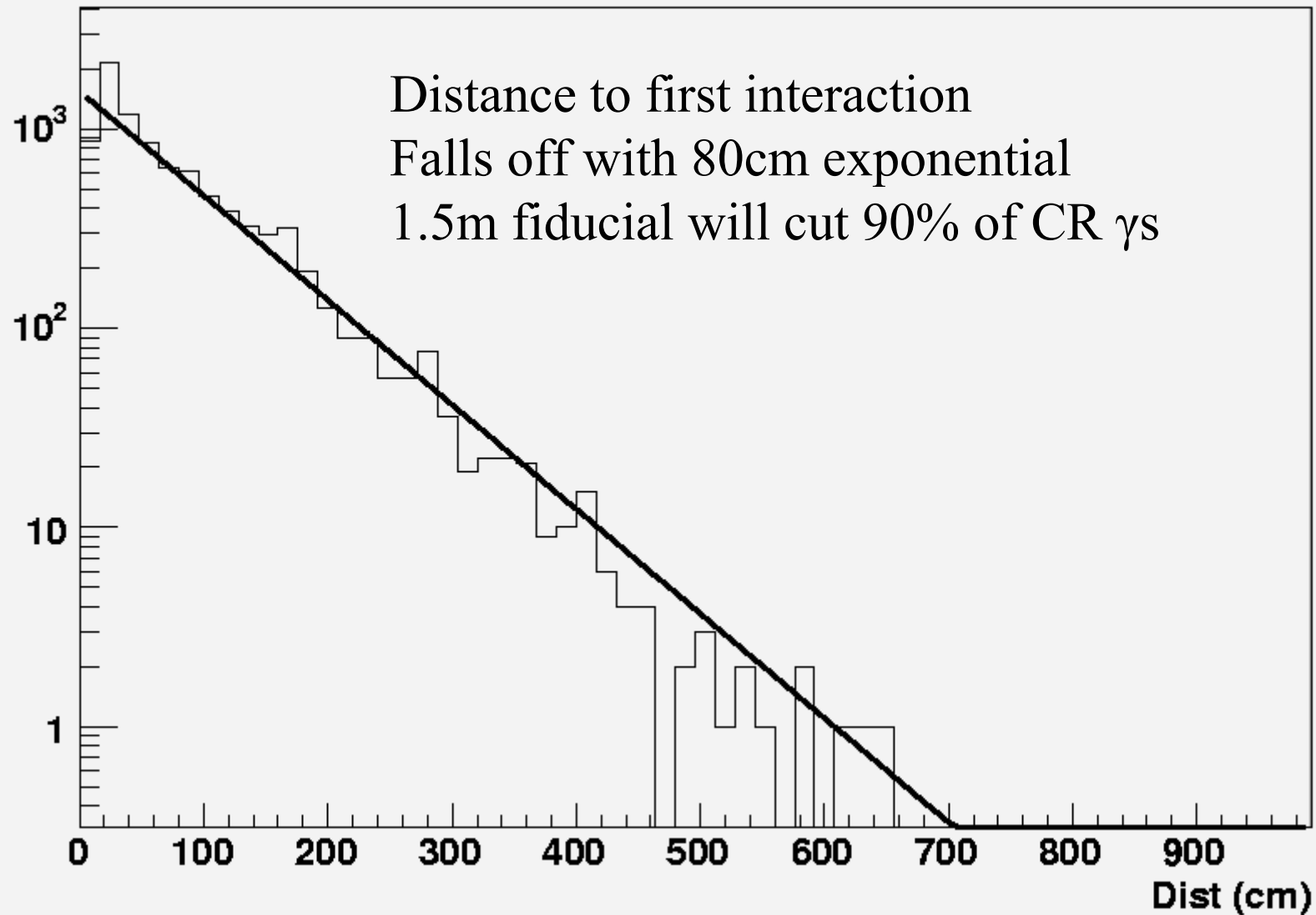
CR gamma Background



CR gamma signals Pulse Height Frequency



Gamma Attenuation



EM Background

- Use Veto region (1.5m)
 - Vetoes 90% of gammas and 100% of e^+e^-
 - $\# \text{photons} / (\#e^+ + \#e^-) = \sim 2$
 - Reduces rate from 300kHz to 20kHz
 - Calorimetry reduces rate $\sim 2x$ (EAS EM higher energy)
 - Not all independent (EAS) should reduce rate further 2? 10x? Corsika could help here.
 - Local (spatially) veto, so still $\sim 90\%$ live mass

Muon Background

- Muon rate in TA detector 500kHz.
- Median energy 4GeV
- About $\frac{1}{2}$ will stop –many Michel electrons
 - Veto for 15 μ s at the end of the track
 - Negligible dead time

Muon Background

- Detector, and any reasonable veto region extremely efficient at detecting muons
- Assume if all data read out muon background (and Michel decays) ==0
- Main problem is horrendous data rate
 - 500kHz muons
 - Generates 300MHz of hits
 - $\sim 3\text{GB/s}$ of data

Muon Veto

An idea:

Use coincidences in outer modules to veto muons.

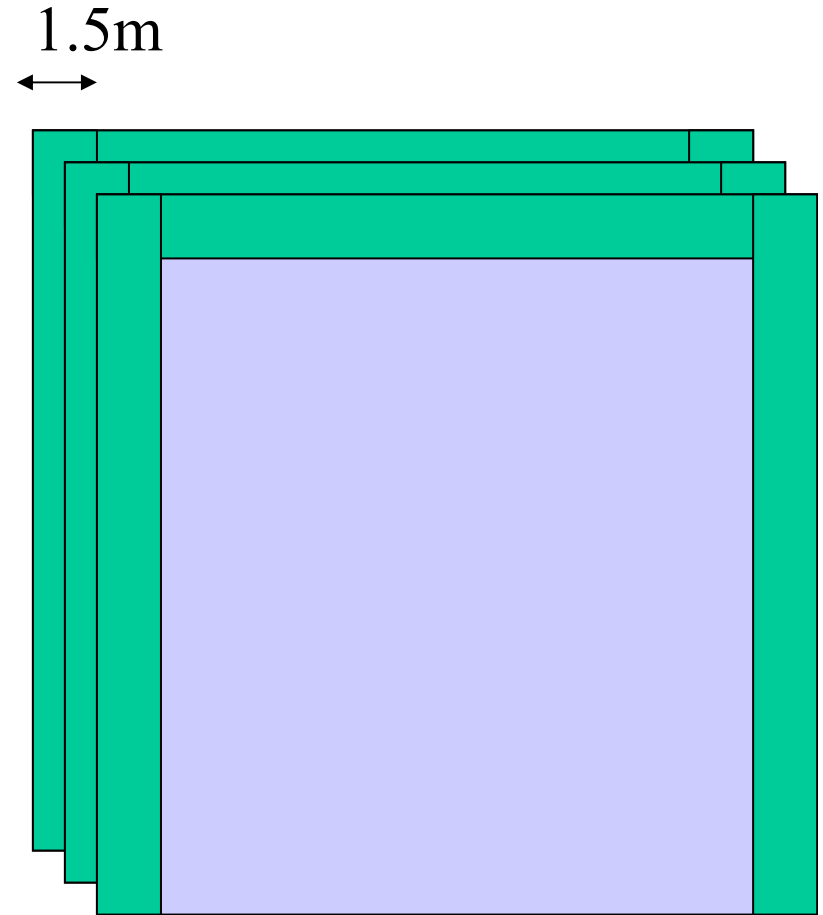
Hardware veto bit sent to a section of the detector, 1/10?

$\sim T/2W = 5/130 = 4\%$ inefficiency 5cm gap= T, W is thickness of veto layer Data Rate $\rightarrow 100\text{MB/s}$

Pipelined and sparsified data from front-end boxes allows decision to only send data for time segments without a veto hit

Possible problem from Michel electrons

Probably need to veto a couple μs , reduces live fraction by $.1 * 2\mu\text{s} / 2\mu\text{s}$, a 10% hit



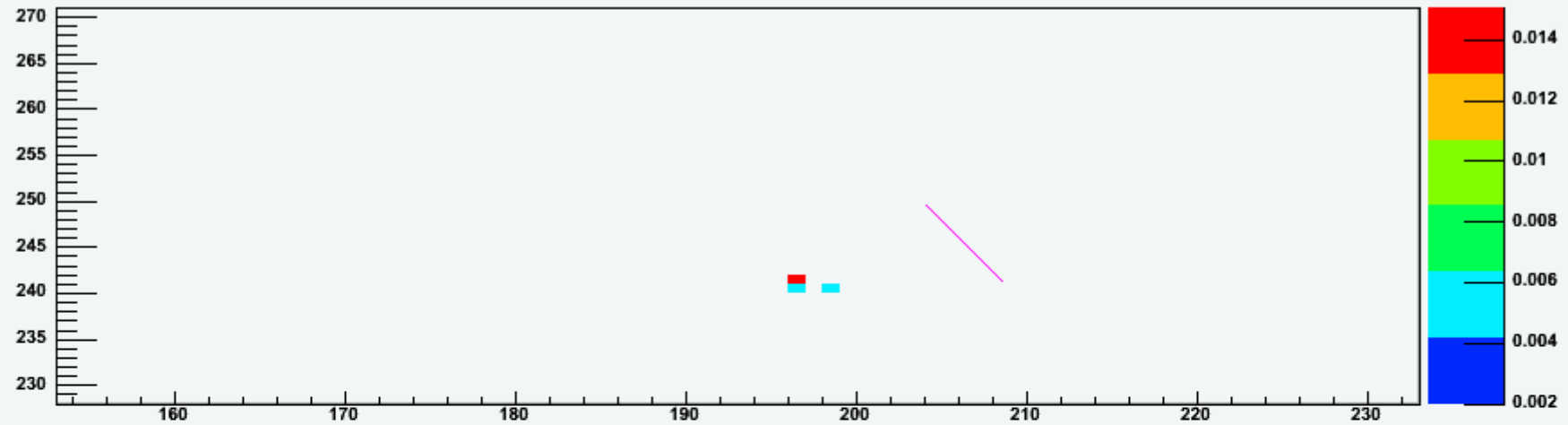
Neutron Background

- Neutron rate 30kHz $E > 100\text{MeV}$

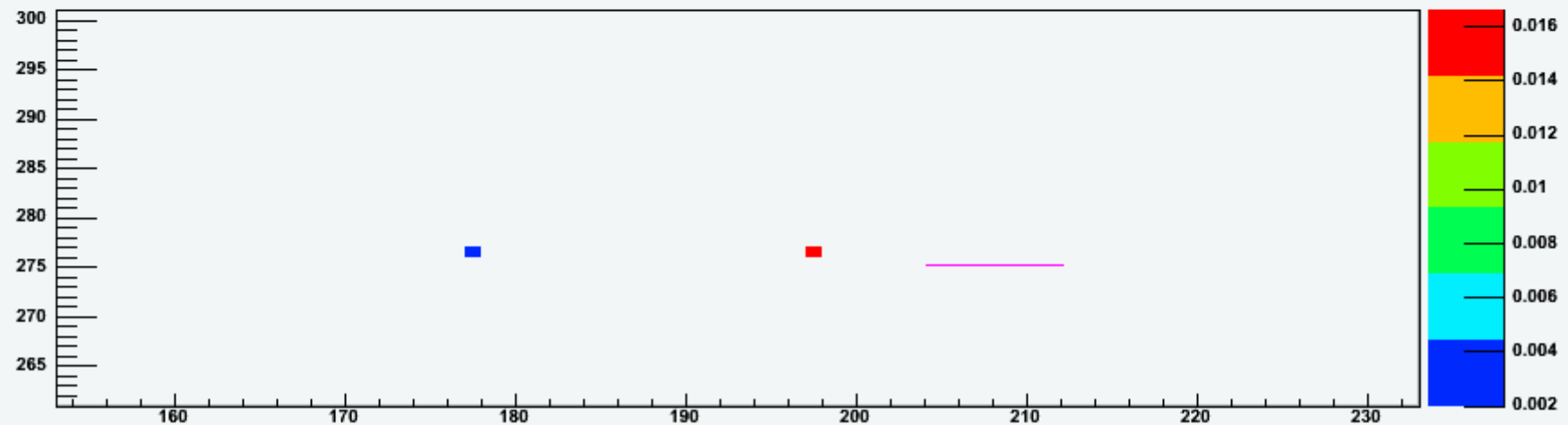
Neutron Background

Event 8 from /usr/scratch/sect2/muallem/TASNNeutrons.root

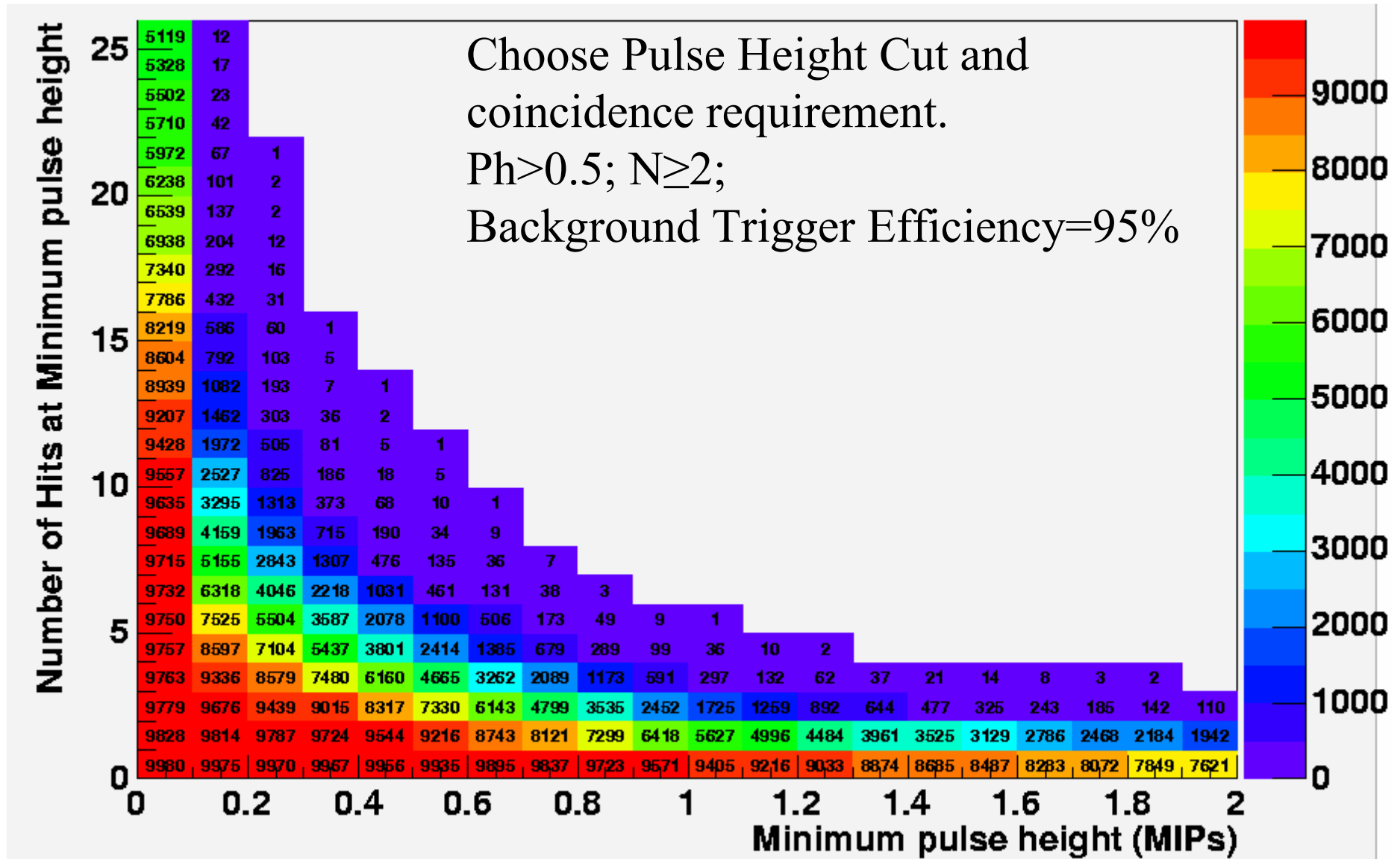
XHitVsPlane



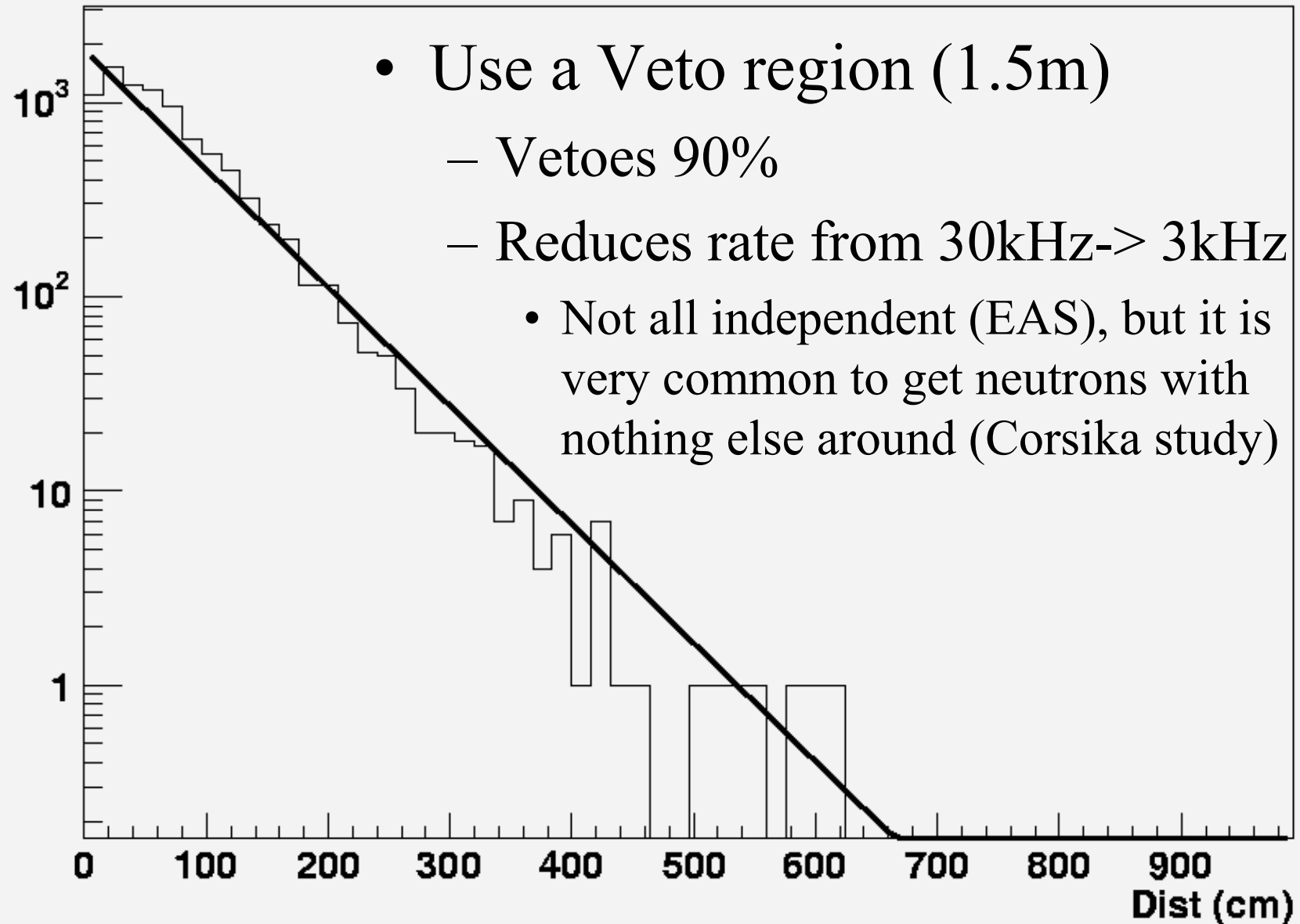
YHitVsPlane



Neutron Pulse Height Frequency



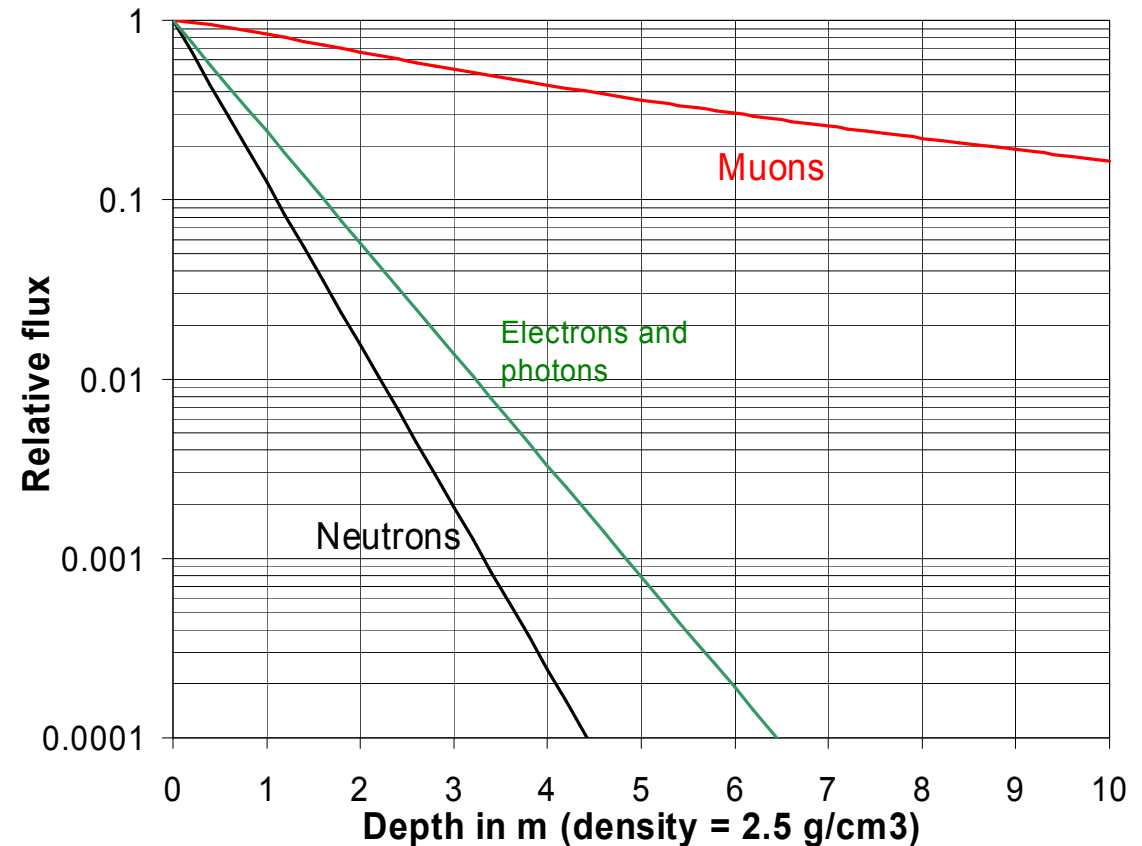
Neutron Attenuation



Effect of overburden

Assume density = 2.5 g/cm²

- Muons:
 - 3m cuts by factor ~2
 - 5m cuts by factor ~3
- Neutrons:
 - 3m cuts by factor ~1000
 - 5m cuts by factor ~3X10⁴
- EM:
 - 3m reduces ~50
 - 5m reduces ~1000



Conclusions

- Expected Signal of approximately 1500 “events” in the first second
- Requires continuous digitization and some hardware muon veto
- Backgrounds:
 - EM 20kHz –Upper limit, probably 10kHz
 - Neutron 3kHz –Upper limit
 - Natural Radioactivity ~ 0 (measurable?)
- FOM $S/\sqrt{Bkd}=10$ sigma minimum, probably more
- With 3m overburden background further reduced
- $\Rightarrow 100 \text{ Hz}$; FOM $\Rightarrow \sim 150$

10MeV Gammas

